# The Gemini Observatory: An Application of High-Performance Networks Tools in Modern Astronomy

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## Why Do Astronomy? Research Objectives

- Understand the Universe Its Origins
  Current State Its "Destination"
- Origins of Life in the Universe
- Connecting Relativity and QCD, (GUTs/TOEs)
- And Other Modest Goals



### Some of Our Tools Gemini North, Mauna Kea Hawaii







#### Typical Gemini Science Missions Primarily Infrared Astronomy

- Evolution and Formation of Elements
- Formation and Evolution of Galaxies
- Black Holes in Galactic Cores
- Stellar Nurseries
- Visualizing Planets Around Other Stars
- Evolution of Planetary Systems



## **Observing Site Requirements**

- Short air column to space
- High percentage of clear skies
- Very dry air from telescope to space
- Low atmospheric turbulence laminar flow
- Low EM "pollution" light or radio
- Usually moderate wind speeds
- Accessibility to scientists



### Where Are These Conditions Found?

• High mountains

On some volcanic tropical islands On some high, arid deserts

- Away from traditional storm tracks
- Smooth up-slope approaching terrain
- Far from major cities

=> Generally places were people don't live

#### Subaru, Mauna Kea Hawaii

#### ALMA, Atacama Desert Chile

Kitt Peak, Arizona

, Central Puerto R



# One Observatory Two Telescopes The Whole Sky



Mauna Kea Hawai'i 13,700 ft

#### Cerro Pachón Chile 9,000 ft





### Where Do the Astronomers Live?

Generally, somewhere far, far away! Where people can breath, plants grow, and shopping malls are handy.

Consider Gemini, for example...



## Gemini's Science Communities A Seven-Nation Partnership





## **Typical Issues**

- Harsh and Remote Environments
- Multi-Site, Multinational Coordination
- Geographically Diverse Communities
- Analysis of Large Data Sets
- Economical Operations
- Effective Communication with the Public



# **Typical Network Approaches**

- Telepresence:
  - Putting the Operators at the Telescope Virtually Sea-level Control Rooms

Putting the Scientist at the Telescope – Virtually Videoconferencing, Access Displays, Participate

- Remote Observing Rooms
- Data Delivery to Scientists and Archives
- Off-site Back Up (!)
- Remote Analysis of Data, Grid Processing
- Network-based Education StarTeachers





802228 (R00352) 2-94

February 1994

North

Pacific

Ocean



### Connecting the Sites The Gemini Backbone

Gemini North & MK Observatories

Gemini Internal Operational Backbone

Gemini South & CTIO, SOAR



#### High-Level External Network Concept Connecting Everybody







#### How Much Bandwidth? One Indicator: Instrument Rates

#### Future Gemini Instruments, Data Rates, and Storage Rates

Instrument Name	Delivery Date	Format	Bits/ Pixel	Bits/ Frame	Frames/ Hour	Transfer Rate (Mbps)	Storage (MB/hr)	Storage (GB/nite)
GMOS-S	2002	4608x6144	16	4.53E+08	18	9.1	1,019	10.2
bHROS	2003	4608x4096	16	3.02E+08	6	10.1	226	2.3
ALTAIR WFS	2003	80x80	16	1.02E+05	8000	20.5	102	1.0
NIFS	2003	2048x2048	32	1.34E+08	24	26.8	403	4.0
GNIRS	2003	1024x1024	32	3.36E+07	50	33.6	210	2.1
FLAMINGOS-2	2005	2048x2048	32	1.34E+08	24	26.8	403	4.0
GSAOI	2005	4096x4096	32	5.37E+08	15	53.7	1,007	10.1
NICI	2005	2048x1024	32	6.71E+07	60	67.1	503	5.0

These figures represent the demand on the internal network, and the requirements of the *external* network when real-time delivery of the images to a remote site is needed.



## **Open Issues**

- Gemini North-South VPN -- Up and Running
- Actually getting the theoretical bandwidth Fine tuning the middleware for long paths "Old" apps like ftp problematic at 250ms RTT (Modern apps ok, e.g. H.323)
- Other service enhancements (more next slide)
- Keeping up with the bandwidth/time profile



# **Coming Enhancements**

- Video-conference enhancements Electronic whiteboards Remote PPT
- VoIP linked to Gemini phone plant
- Remote telescope "viewing rooms"
- Explore moving data to archives on line
- Establish Access Grid Node
- Migration to IPv6
- Enhanced Gemini South Bandwidth, as required



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